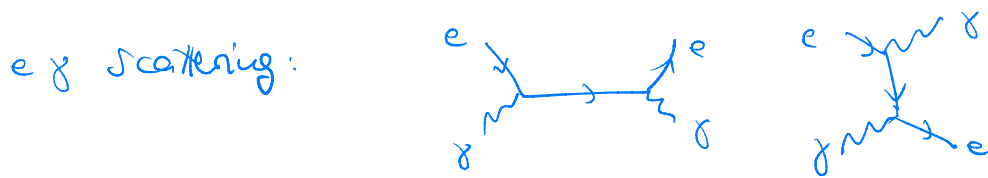


5.2 Photon decoupling

also called "photon last scattering".



for $energy \ll m_e$ the cross section is given by
the Thomson cross section

$$\sigma_T \approx \frac{8\pi}{3} \frac{\alpha^2}{m_e^2} \approx 6.7 \times 10^{-25} \text{ cm}^2$$

N.B. The cross section for γp scattering is much smaller!
Thermal equilibrium is maintained by Coulomb interaction
of e and p .

Boltzmann eq: $\partial_t f_\gamma = f_\gamma n_e \sigma_T \Rightarrow$

mean free time of a photon

$$\tau_\gamma = \frac{1}{n_e \sigma_T}$$

$$\text{At } T_{\text{rec}}: \quad n_e = 0.67 \times 1.8 \times 10^{-11} \eta_{10}^3 (0.31 \text{ eV})^3 \\ \simeq 2.1 \times 10^{-12} (\text{eV})^3$$

$$\sigma_T = 1.7 \times 10^{-15} (\text{eV})^{-2}$$

$$\tau_\gamma = 2.7 \cdot 10^{26} \text{ eV}^{-1}$$

$$t_1 = 6.7 \times 10^{-22} \text{ MeV}^{-1} = 6.7 \times 10^{-16} \text{ eV}^{-1}$$

$$\tau_\gamma = 1.8 \times 10^{-11} \text{ s} \ll H_{\text{rec}}^{-1}$$

so then the photons are still in kin. equilibrium with electrons.

Therefore photons decouple later than recombination.

Since n_e is then no longer in equilibrium, one needs to solve Boltzmann equations to determine the decoupling time.

$$\text{result:} \quad T_{\text{dec}} \simeq 0.26 \text{ eV} \quad [\text{Gorbunov Rubakov 6.2}]$$

6 Electroweak epoch

6.1 Symmetry restoration

In the standard model of particle physics all fundamental constituents are quanta of some fields. Historically, the first example was the quantum of the electromagnetic field.

matter ($\text{spin } \frac{1}{2}$)

quarks, leptons

force carrier/gauge bosons ($\text{spin } 1$)

$\gamma, W^{\pm}, Z, g$

weak interactions

gluons, strong interactions

Higgs field ($\text{spin } 0$)

The quarks, charged leptons and weak gauge bosons get their masses from interaction with the Higgs field

Yukawa interaction for leptons

$$\mathcal{L}_{\text{Yukawa}} = -y \bar{L} \varphi l + \text{c.c.}$$

$\bar{L}, l$ $\text{spin } \frac{1}{2}$ fields

For one fermion family (we have actually three)

$L = \begin{pmatrix} \nu \\ e \end{pmatrix}_L$ left-handed doublet

$l = e_R$ right-handed electron

$$\varphi = \begin{pmatrix} \varphi_1 \\ \varphi_2 \end{pmatrix}$$

y : Yukawa coupling

In the vacuum the Higgs field has a non-zero constant value $\langle \varphi \rangle = \begin{pmatrix} 0 \\ v/\sqrt{2} \end{pmatrix}$, $v \simeq 246 \text{ GeV}$

This gives a mass term

$$\mathcal{L}_{\text{mass}} = -\frac{y}{\sqrt{2}} v \bar{e}_L e_R + \text{c.c.}$$

for the electron, while neutrinos remain massless.

This term is not invariant under the SM gauge transformations, $\bar{L} \rightarrow \bar{L} U^\dagger$, $e_R \rightarrow V e_R$ with $U \in SU(2) \times U(1)$, $V \in U(1)$. That's why it's called electroweak symmetry breaking.

N.B. Most of the mass of SM matter in our universe is due to binding effects of the strong interaction.

Example: the proton contains 3 quarks and, held together by gluons

quark masses $m_u, m_d \lesssim 10 \text{ MeV}$

gluons are massless

But: proton mass $m_p \simeq 930 \text{ MeV}$

At $T \gg 250 \text{ MeV}$ the running coupling of QCD $\alpha_s(T)$ becomes small. Then quarks and gluons are no longer confined to hadrons and can be approximately viewed as weakly interacting particles, just like leptons and photons.

high temperatures

Suppose we are at temperatures T bigger than the mass of the heaviest SM particle, the top quark with $m_{\text{top}} \approx 176 \text{ GeV}$

Then the pressure is given by its massless limit plus some small correction due to the mass

$$P = g T^4 \begin{cases} \frac{\pi^2}{90} - \frac{m^2}{24 T^2} + \dots & \text{bosons} \\ \frac{7}{8} \frac{\pi^2}{90} - \frac{m^2}{48 T^2} + \dots & \text{fermions} \end{cases}$$

The m^2 -correction is negative because the mass reduces the momentum of the particles for given energy, giving less pressure.

The Higgs field is a dynamical field, so that its value can change.

$$P = c_0 T^4 - c_2 T^2 |\langle \varphi \rangle|^2 + \dots$$

with $c_0, c_2 > 0$

regions with smaller $|\langle \varphi \rangle|$ have larger pressure and are favored $\Rightarrow$

$$\langle \varphi \rangle = 0 \quad \text{at high } T$$

electroweak symmetry restoration

Thus all particles become massless at $T \gtrsim m_{\text{top}}$

A more precise computation shows that this

happens for $T \gtrsim 130 \text{ GeV}$