

2.4 Lepton era

Coupled $1 \text{ MeV} \lesssim T \lesssim 100 \text{ MeV}$

quarks & gluons are tightly bound into hadrons

lightest hadrons: pions, $m_\pi \sim 140 \text{ MeV} \ll T$ non-relativistic

same for proton ($m_p = 938 \text{ MeV}$), neutrons ($m_n = 939 \text{ MeV}$)

muons ($m_\mu \approx 113 \text{ MeV}$) and τ -leptons ($m_\tau \approx 1.78 \text{ GeV}$)

The only relativistic particles left are

$\gamma, e^\pm$, (anti-)neutrinos $\nu_\alpha, \bar{\nu}_\alpha$ ($\alpha = e, \mu, \tau$)

Conserved are el. charge Q , baryon number B , and

(for $T \gtrsim 1 \text{ MeV}$) 3 lepton flavor numbers L_α

particle chemical potentials:

$$\mu_{\pi^0} = 0, \quad \mu_{\pi^\pm} = \pm \mu_Q$$

$$\mu_e = -\mu_Q + \mu_{L_e} \quad \mu_\mu = -\mu_Q + \mu_{L_\mu}, \dots$$

$$\mu_p = \mu_Q + \mu_B \quad \mu_n = \mu_B$$

The values of Q , B and L_α together with

$$\frac{d(\gamma a^3)}{dt}$$

determine the functions $T(t), \mu_Q(t), \mu_B(t), \mu_{L_\alpha}(t)$

$$Q = 0, \quad \frac{n_B}{s} \sim 10^{-10}$$

The lepton numbers are very weakly constrained because it is very difficult to measure densities of neutrinos & antineutrinos.

From observations $n_{L\alpha} \gg n_3$ cannot be excluded.

On the other hand, most theories which can explain

the origin of the value of n_3 also predict $n_{L\alpha} \sim n_3$

For $T \gg m_i$, then $n_i \sim n_{\bar{i}} \sim n_\gamma \sim T^3$

When T drops below m_i , most particles and anti-particles annihilate.

Only the excess due to the asymmetries survives.

Assumption: some asymmetry is carried by one species i

only $\eta = \frac{\Delta n_i}{s} = \text{const}$ with $\Delta n_i = n_i - n_{\bar{i}}$

for nonrelativistic particles we have

$$n_i = g_i \left(\frac{m_i T}{2\pi} \right)^{3/2} e^{(\mu_i - m_i)/T}$$

$$\mu_{\bar{i}} = -\mu_i$$

$$s = g_* \frac{2\pi^2}{45} T^3$$

$$\frac{n_i n_{\bar{i}}}{s^2} = g_i^{-2} \frac{45^2}{2^5 \pi^7} T^{-3} g_i^2 \left(\frac{m_i T}{2\pi} \right)^3 e^{-2m_i/T}$$

$$= c^2 \frac{g_i^2}{g_*^2} \left(\frac{m_i}{T} \right)^3 e^{-2m_i/T}$$

with $c^2 = \frac{45^2}{2^5 \pi^7} \approx 0,021$

$$n_{\bar{i}} = n_i - \gamma s \Rightarrow$$

$$\frac{n_i n_{\bar{i}}}{s^2} = \frac{n_i}{s} \left(\frac{n_i}{s} - \gamma \right) = \left(\frac{n_i}{s} \right)^2 - \gamma \frac{n_i}{s}$$

$$\frac{n_i}{s} = \frac{\gamma}{2} + \left[\frac{\gamma^2}{4} + c^2 \frac{g_i^2}{g_*^2} \left(\frac{m_i}{T} \right)^3 e^{-2m_i/T} \right]^{1/2}$$

exchange $i \leftrightarrow \bar{i} \Leftrightarrow \gamma \leftrightarrow -\gamma$:

$$\frac{n_{\bar{i}}}{s} = -\frac{\gamma}{2} + \left[\frac{\gamma^2}{4} + c^2 \frac{g_i^2}{g_*^2} \left(\frac{m_i}{T} \right)^3 e^{-2m_i/T} \right]^{1/2}$$

number of particle - antiparticle pairs negligible compared to the particle excess number when

$$(*) \quad \frac{\gamma}{2} \gg c \frac{g_i}{g_*} \left(\frac{m_i}{T} \right)^{3/2} e^{-m_i/T}$$

rough estimate: $(*) \Leftrightarrow \frac{m_i}{T} \gg \ln \frac{2c}{\gamma}$

for $\gamma = 10^{-10}$: $\ln \frac{2c}{\gamma} \approx 22$, $T \ll \frac{m_i}{22}$

$$m_p = 938 \text{ MeV}$$

$$\# \text{ antibaryons} \ll \# \text{ baryons for } T \lesssim 43 \text{ MeV}$$

$$m_e = 511 \text{ keV}$$

$$\# e^+ \ll \# e^- \text{ for } T \lesssim 23 \text{ keV}$$

estimate chemical potentials

$$\text{for } T \ll 100 \text{ MeV}, \quad m_p, m_\tau \gg T \text{ while } m_\nu, m_{\nu_e} \ll T$$

$$\Rightarrow L_\mu, L_\tau \text{ mostly carried by } \nu_\mu, \nu_\tau$$

$$\text{problem H6.1: } \Delta n_i = \frac{\mu_i T^2}{6} \Rightarrow$$

$$\frac{\mu_{L\alpha}}{T} \sim \frac{n_{L\alpha}}{s} \quad \text{for } \alpha = \mu, \tau$$

The contribution of e^- to L_e cannot be neglected, because the electrons must compensate the electric charge carried by the protons.

Relativistic electrons.

$$\left. \begin{aligned} \Delta n_{e^-} &= 2 \frac{\mu_{e^-} T^2}{6} \\ \Delta n_{e^-} &\sim n_B \end{aligned} \right\} \Rightarrow \frac{\mu_{e^-}}{T} \sim \frac{n_B}{s}$$

$$\text{Furthermore, } \mu_{e^-} = \mu_{L_e} - \mu_a$$

$$n_{Le} \simeq 2 \Delta n_{e^-} + \Delta n_{\nu_e} \simeq (2\mu_{e^-} + \mu_{\nu_e}) \frac{\hat{T}^2}{6}$$

$$= (2\mu_{Le} - 2\mu_Q + \mu_{Le}) \frac{\hat{T}^2}{6} = (3\mu_{Le} - 2\mu_Q) \frac{\hat{T}^2}{6}$$

Thus we see that n_{Le} can be much larger than $n_{e^-} = n_B$

Furthermore, we see that for relativistic particles $\frac{\mu_i}{T} \simeq \text{const}$

Now consider p, n .

For $T \ll 45 \text{ MeV}$: $\Delta n_p \simeq n_p$, $\Delta n_n \simeq n_n$

$$\gamma_B = \frac{n_B}{s} \simeq \frac{n_p + n_n}{s} \simeq \frac{n_p}{s} \left(1 + \frac{n_n}{n_p}\right)$$

$$g_p = 2 \Rightarrow$$

$$n_p = 2 \left(\frac{m_p T}{2\pi} \right)^{3/2} e^{(\mu_p - m_p)/T}$$

$$s = g_* \frac{2\pi^2}{45} T^3$$

assume $n_n \simeq n_p$ (will be checked later)

$$\gamma_B \simeq \frac{2}{g_*} \underbrace{2^{-5/2} \pi^{-3/2} 45}_{= c} \left(\frac{m_p}{T} \right)^{3/2} e^{(\mu_p - m_p)/T} \cdot 2$$

$$\ln \gamma_B \simeq \cancel{\frac{\mu_p - m_p}{T}} + \ln\left(\frac{4c}{g_*}\right) + \frac{3}{2} \ln\left(\frac{m_p}{T}\right)$$

$$\cancel{\frac{m_p - \mu_p}{T}} = \ln\left(\frac{4c}{\gamma_B} \left[\frac{m_p}{T}\right]^{3/2}\right)$$

$$T = 40 \text{ MeV}: \quad \ln(\dots) = 25 \quad \Rightarrow \mu_p \simeq -100 \text{ MeV}$$

$$T = 1 \text{ MeV} : \quad \ln(\dots) = 30.4$$

$$\mu_p = m_p - 30.4 \text{ MeV} = 908 \text{ MeV}$$

entropy carried by protons:

We had $s_i = n_i \left(\frac{m_i - \mu_i}{T} + \frac{5}{2} \right)$ for NR particles

$$\frac{s_p}{s} \simeq \eta_3 \left\{ \ln \left(\frac{4c}{\eta_3} \left[\frac{m_p}{T} \right]^{3/2} \right) + \frac{5}{2} \right\} \sim 10^{-9}$$

tiny fraction of the total entropy!

now estimate

$$\frac{n_n}{n_p} = \exp \left(\frac{1}{T} [\mu_n - m_n - \mu_p + m_p] \right)$$

$$\mu_n = \mu_B, \quad \mu_p = \mu_B + \mu_Q, \quad \mu_n - \mu_p = \mu_Q$$

$$\mu_Q \sim \mu_{Le} \sim \eta_{Le} T \ll T$$

$$\Delta m := m_n - m_p = 1.29 \text{ MeV}$$

$$\frac{n_n}{n_p} \simeq \exp \left(-\frac{\Delta m}{T} \right) \simeq 1$$

for $T \gtrsim 1 \text{ MeV}$ ok