

that was 'multiplicative renormalization'.

now: counterterms

$$\delta = \frac{1}{2}(\partial_\mu f_0) \partial^\mu f_0 - \frac{1}{2} m_0^2 f_0^2 - \frac{\lambda_0}{4!} f_0^4$$

$$f_0 = \sqrt{Z_f} \phi, \quad m_0^2 = m^2 + \delta m^2, \quad \lambda_0 = \frac{Z_f}{\lambda^2} \lambda$$

(index '0' omitted from now on)

$$= \frac{1}{2}(\partial_\mu \phi) \partial^\mu \phi - \frac{1}{2} m^2 \phi^2 + \delta_{\text{int}}$$

$$\delta_{\text{int}} = -\frac{\lambda}{4!} \underbrace{Z_f \phi^4}_{O(\lambda^2)} - \frac{1}{2} \underbrace{\delta m^2 Z_f \phi^2}_{O(\lambda^2)} + \frac{1}{2} \underbrace{(Z_f - 1)(\lambda \phi)^2 - m^2 \phi^2}_{O(\lambda^2)}$$

$$= -\frac{\lambda}{4!} \phi^4 - \frac{1}{2} \left(\frac{\delta m^2}{\lambda} \right) \phi^2 + O(\lambda^2)$$

(end of diagram)



→ this way, the (mult.) r.m. can be incorporated into our diagrammatic expansion:

$$\ln Z_{\text{int}} = 300 + 0 + O(\lambda^2)$$

$$G_2 = G_0 + 12 \underbrace{\square}_{\substack{O(2+\text{int})}} + 2 \underbrace{\star}_{\substack{\text{int}}} + O(\lambda^2)$$

$$\Pi = -12 \underbrace{\square}_{\substack{O(2)}} - 2 \underbrace{\star}_{\substack{\text{int}}} + O(\lambda^2)$$

$$\begin{aligned} &= -12 \left(-\frac{\lambda}{4!} \right) \underbrace{\sum_n G_0(n)}_{I^{\text{inc}} + I^{\text{int}}} - 2 \left(-\frac{1}{2} \delta m^2 \right) \underbrace{-\Pi_{\text{ext}}(m)}_{\lambda} = 12 \left(-\frac{\lambda}{4!} \right) I^{\text{inc}} \\ &= -12 \left(-\frac{\lambda}{4!} \right) I^{\text{int}} \end{aligned} \quad (96)$$

now, we can finally

• evaluate $\ln Z^{(n)}$

$$\ln Z^{(n)} = 300 + 0$$

$$= 3 \left(-\frac{\lambda}{4!} \right) \beta V (I^{\text{inc}} + I^{\text{int}})^2 + \left(-\frac{1}{2} \right) 12 \left(-\frac{\lambda}{4!} \right) I^{\text{inc}} \beta V (I^{\text{inc}} + I^{\text{int}})$$

$$= 3 \left(-\frac{\lambda}{4!} \right) \beta V [(I^{\text{int}})^2 - (I^{\text{inc}})^2] \quad (97)$$

→ in (97), the terms vac-mat have cancelled !!!

still, $\ln Z^{(1)}$ contains divergences: $(I^{vac})^2$, but is not physical.

physical quantities $F = -T \partial_T \ln Z$, $p = T \partial_T \ln Z$?
 \downarrow
 $p = p_{mat} + p^{(1)} + \dots$

$$\begin{aligned} p^{(1)} &= T \partial_T \ln Z^{(1)} \\ &= 3 \left(-\frac{\lambda}{4!} \right) [(I^{mat})^2 - (I^{vac})^2] \\ &= p_{mat}^{(1)} + p_{vac}^{(1)} \quad (\text{splits only because no vac-mat in } \ln Z^{(1)}) \end{aligned}$$

one measures only differences (in energy, pressure, ...)

→ can choose to normalize $p_{vac} = 0$

$$\Rightarrow p_{phys}^{(1)} = p_{mat}^{(1)} = 3 \left(-\frac{\lambda}{4!} \right) \underbrace{\left(\frac{T^2}{2\pi^2} P(3) h_3(\zeta) \right)^2}_{m=0} \quad (\zeta = \frac{T}{\Lambda}) \approx \frac{T^4}{144} (h_0(\frac{T}{\Lambda})) \quad (98)$$

• evaluate $\ln Z^{(2)}$

$$\begin{aligned} \ln Z^{(2)} &= 36 \ 000 + 12 \ 000 + \text{ct's} \quad \checkmark \text{containing} \\ &= 36 \beta V \left(\frac{\pi^6}{-12} \right)^2 \sum_{Q_1, Q_2} [Q_1 + Q_2] G_0(Q_1) G_0(Q_2) \\ &\quad + 12 \beta V \left(-\frac{\lambda}{4!} \right)^2 \sum_{Q_1, \dots, Q_4} [Q_1 + \dots + Q_4] G_0(Q_1) \dots G_0(Q_4) + \text{ct's} \end{aligned}$$

... heading for next term in pressure $\boxed{\alpha_{m=0}} \dots$

(could now proceed to calc above int's, normalize (add ct's), and then be surprised that a (new type of) divergence occurs. So, to make the story shorter, investigate cycle first!)

look at IR (infrared; small momentum) behavior

$$2^{\text{nd}} \text{ day}, \quad n_1 = n_2 = n_3 (= n_4) = 0 \quad \text{from } (T \sum_n)^3$$

$$\sim \overleftarrow{T^3} \int d^3 p_1 \dots d^3 p_4 \frac{1}{p_1^2} \dots \frac{1}{p_4^2} \delta(p_1 + \dots + p_4)$$

$$= T^3 \int d^3 p_1 \dots d^3 p_3 \frac{1}{p_1^2} \frac{1}{p_2^2} \frac{1}{p_3^2} \frac{1}{(p_1 + p_2 + p_3)^2}$$

$$\sim T^3 \int dp_1 \dots dp_3 \quad 1 \cdot 1 \cdot 1 \cdot \frac{1}{(p_1 p_2 p_3)^2}$$

$$\sim T^3 dp \quad \text{IR convergent} \quad \checkmark \text{ no diverge.}$$

$$\text{1st drag} \sim T \overbrace{\int d^3 p_1 d^3 p_2}^{\text{from } T \Sigma} \frac{1}{p_1^2} \frac{1}{p_2^2} \delta(p_1, p_2)$$

$$= T \int d^3 p_1 \frac{1}{p_1^2} \frac{1}{p_1^2}$$

$$\sim T \int d\vec{p}_1 1 - \frac{1}{p_1^2} \quad \text{IR divergent!}$$

now $T \neq 0$ - effect! vanishes at $T \rightarrow 0$

((physical reason for div: expanding with massless propagator;

know that $\Pi_i^{ren} = \frac{1}{4!} T^2$ is dynamically generated mass term))

IR div. gets worse in higher orders:



$$\sim \rho V (\Pi_i)^N T \int d^3 p \left(\frac{1}{p^2}\right)^N \sim \rho V (\Pi_i)^N T \int d\vec{p} \left(\frac{1}{p^2}\right)^{N+1}$$

simple comb. factor: $\frac{1}{N!} (3! \cdot 2)^N \frac{(N-1)!}{2}$ ((check: $N=2, \frac{1}{2!} (3! \cdot 2)^2 \frac{1!}{2} = 36$)

e.g. ↑ ↑ ↑ ↑ # of ways to order
 connect connect to loop 1..N on circle
 ... O ... 2 2 2 2

due to this simple structure, this class of drag can actually be summed to all orders:

$$\begin{aligned}
 0O + \text{diag} + \dots &= \sum_{n=2}^{\infty} \frac{(3! \cdot 2)^n}{2^n} \rho V \left(\frac{\Pi_i}{\pi^2}\right)^n \sum_k (G_0(k))^n \\
 &= \frac{1}{2} \rho V \sum_n \sum_{k=2}^{\infty} \frac{(-)^n}{n} (\Pi_i G_0(k))^n \\
 &= \frac{1}{2} \rho V \sum_n \underbrace{\left(\sum_{k=1}^{\infty} \frac{(-)^n}{n} (\Pi_i G_0)^n + \Pi_i G_0 \right)}_{= -\ln(1 + \Pi_i G_0)} \quad (98a)
 \end{aligned}$$

'ring diagrams';

include mass counterterm $\star \Rightarrow \Pi_i \rightarrow \Pi_i^{ren} \left(= -2 \left(-\frac{1}{4!} \right) I_{loop}^{\text{mat}} / m_0 = \frac{1}{4!} T^2 \right)$
(proof: see hw)

\Rightarrow using $\Pi_{\text{ren}} = \frac{\lambda}{4!} T^2 \quad (\text{m=0})$,

ring diag's $= -\frac{1}{2} \rho V \sum \left(\ln \left(1 + \frac{\frac{\lambda}{4!} T^2}{\omega^2 + \epsilon^2} \right) - \frac{\frac{\lambda}{4!} T^2}{\omega^2 + \epsilon^2} \right)$

$= -\frac{1}{2} \rho V T \int \frac{d\omega}{(2\pi)^3} \left(\ln \left(1 + \frac{\frac{\lambda}{4!} T^2}{\epsilon^2} \right) - \frac{\frac{\lambda}{4!} T^2}{\epsilon^2} \right) \quad + \text{massless modes}$

$\xrightarrow{\text{static (m=0) modes only}}$ $- \frac{1}{2} \rho V T \sqrt{\frac{\lambda}{4!} T^2} \frac{1}{2\pi^2} \int d\epsilon \epsilon^2 \left(\ln \left(1 + \frac{1}{\epsilon^2} \right) - \frac{1}{\epsilon^2} \right) \quad \begin{matrix} \text{(into-modes;} \\ \text{one higher corr. m=1)} \end{matrix}$

finite integral ✓ makes UV well behaved

(in particular, IR finite !!!)

$= \frac{\beta V}{12\pi} \left(\frac{\lambda}{4!} \right)^{3/2} T^4 \quad (996)$

\Rightarrow the 'NNLO' of λT (and F, p, \dots) is actually of order $\lambda^{3/2}$, not λ^2 like in naive (no resummations) pert. expansion!

collecting from (65), (98) and (996),
we get for the pressure of massless $\frac{\lambda}{4!} T^4$ - theory

$$P_{\text{phys}} = T^4 \left(\frac{\pi^2}{90} - \frac{1}{48} \frac{\lambda}{4!} + \frac{1}{120} \left(\frac{\lambda}{4!} \right)^{3/2} + O(\lambda^2) \right) \quad (100)$$

... higher orders: see literature

e.g. [Parvani/Singh, PRD 51 (1995) 4518]

and ref's therein

$\sim \lambda^{5/2}$

convergence ... resummation ...